

Heat Transfer: Radiation

- All things radiate electromagnetic energy

$$\rightarrow I_{\text{emit}} = Q/t = eA\sigma T^4$$

» e = emissivity (between 0 and 1)

■ perfect “black body” has $e=1$

» T is temperature of object in Kelvin

» σ = Stefan-Boltzmann constant = $5.67 \times 10^{-8} \text{ J/s-m}^2\text{-K}^4$

→ No “medium” required



DEMO

- All things absorb energy from surroundings

$$\rightarrow I_{\text{absorb}} = eA\sigma T_0^4$$

» T_0 is temperature of surroundings in Kelvin

» good emitters (e close to 1) are also good absorbers